

# **Turbulent transport regimes and the SOL heat flux width**

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*Presented at the 2014 US TTF Workshop, April 22-25, 2014, San Antonio, Texas.*

*Work supported by US DOE grant DE-FG02-97ER54392 .*

# Introduction

- Need to understand the responsible mechanisms and resulting scaling of the SOL heat flux width  $\lambda_q$ .
  - Critical for viable operation of future machines, e.g.  $\lambda_q$  in ITER
- Simulation and theory using reduced edge/SOL turbulence models<sup>1-5</sup> show some agreement ( $\lambda_q$  magnitude and scaling) with experiment.
  - Overarching qualitative conceptual understanding of regimes and results is still lacking  $\Leftarrow$  **the main goal of this work**
- The heuristic drift model<sup>6</sup> (associated with neoclassical effects<sup>7</sup>) has also been very successful in explaining experimental observations.
  - Here we also speculate on the relationship between the SOL turbulence and drift-based mechanisms.

[1] D.A. Russell et al., Phys. Plasmas **19**, 082311 (2012); and this meeting.

[2] J.R. Myra et al., Phys. Plasmas **18**, 012305 (2011).

[3] F. D. Halpern et al., Nucl. Fusion **53**, 122001 (2013) .

[4] F. Militello et al., Plasma Phys. Control. Fusion **55**, 074010 (2013).

[5] J. W. Connor et al., Nucl. Fusion **39**, 169 (1999).

[6] R.J. Goldston, Nucl. Fusion **52**, 013009 (2012).

[7] C.S. Chang et al., FY-2010 JRT and follow-on work (unpublished)

## Turbulence provides a mechanism for sustaining the SOL width $\lambda_q$

- Heat flux balance in the SOL
  - $\nabla_{\parallel} q_{\parallel} = -\nabla_{\perp} \cdot q_{\perp}$
  - $q_{\perp} \sim D_{\text{turb}} p / \lambda_p$  where  $D_{\text{turb}} \sim \gamma \langle \tilde{v}_x^2 \rangle / |\omega|^2$
  - perpendicular scale length of pressure is  $\lambda_p$ 
    - $1/\lambda_p$  drives the turbulence
  - parallel heat flux  $q_{\parallel} = g n T c_s$ 
    - $g$  is regime-dependent factor
  - parallel scale is  $L_{\parallel} = qR$
- Instabilities of interest
  - **curvature** driven (ideal, resistive, resistive X-pt = RX)
  - Kelvin-Helmholtz (**KH**)
  - collisional drift-wave (**DW**)
  - others?

$$\lambda_q = \frac{qR\gamma \langle \tilde{v}_x^2 \rangle}{g c_s |\omega|^2 \lambda_p}$$

need specific  
inputs for  $\gamma$ ,  
saturation ...

## Fluctuation amplitude $\tilde{v}_x$ is determined by saturation: several regimes are possible

- **Wave-breaking**  $\frac{k_x \tilde{v}_{Ex}}{\omega} \sim 1$ 
  - equivalent to equating the perturbed and equilibrium pressure gradients (pressure-convective saturation)

- Shear flow generation from **Reynolds** stress

$$\gamma = v'_{Ey} \equiv \frac{v_{Ey}}{\lambda_E} = \frac{k_x k_y}{v} \frac{\langle \tilde{\Phi}^2 \rangle}{\lambda_E^2}$$

- here  $\lambda_E$  = scale length of the radial electric field
- and  $v$  = zonal flow dissipation rate
- beats wave-breaking when  $k_x k_y \lambda_E^2 v < \gamma$  or for global modes when  $v < \gamma$
- **Mean flow** suppression (not really a saturation mechanism)
  - an important case is H-mode where we estimate that  $E \times B$  and diamagnetic flows balance

$$\gamma = v'_{di} \equiv \frac{c_s \rho_s}{\lambda_p^2}$$

## Relevant wave-number estimates depend on the regime and type of instability

- **Quasi-local** limit  $k_y \lambda_p \gg 1$ 
  - Conventional resistive modes typically require a high  $k_y \sim k_\eta \equiv \left( \frac{\Omega_e}{v_e} \right)^{1/2} \frac{\lambda_p^{1/4}}{qR^{3/4} \rho_s^{1/2}}$ 

$$\text{from } \omega_\eta \gamma_{\text{mhd}} \sim \omega_a^2$$
  - FLR-mitigated ballooning mode spectrum can peak where  $\gamma \sim \omega_{*i} = k_y c_s \rho_s / \lambda_p$
  - For quasi-local modes, estimate  $k_x^2 \sim k_y \lambda_p$ 
    - obtained from parabolic expansion of a generic eigenmode equation about the point of maximum growth.
- **Non-local** (global) modes  $k \lambda_p \sim 1$ 
  - For these modes the radial eigenfunction overlaps the bulk of the driving gradient
  - Estimate  $k_x \sim k_y \sim 1/\lambda_p$
- **Barrier limited** non-local modes  $k_x \sim \pi/L_x$ 
  - Rapid changes in geometry or plasma profiles near the separatrix can radially confine low  $k_y$  modes to a smaller scale than they would otherwise have.
  - e.g.: electron adiabaticity barrier limits extent of interchange modes; sheaths may do the same: RX modes are limited by the width of the X-pt shear layer

## The instability drive may, or may not, be local to the SOL

- “**Compact**” modes
  - When the driving gradients are in the SOL we estimate
$$\lambda_p \sim \lambda_q$$
  - This is frequently the case for quasi-local instabilities
- “**Distributed**” modes
  - When the driving gradients are in the edge pedestal but large scale convective motions cause this turbulence to govern the SOL width<sup>1,2</sup> then  $\lambda_p$  and  $\lambda_q$  are independent parameters.
  - Here we regard  $\lambda_p =$  driving gradient = an *input*
    - later we discuss a rough estimate for  $\lambda_p$  in an H-mode.
  - $\lambda_q =$  responding gradient = an *output*
  - Note that this distributed mode paradigm connects the pedestal properties to the SOL heat flux width.

## Combining these leads to many possible scalings for $\lambda_q$

- 3+ types of instabilities
  - curvature-driven ideal or RX (a low k version of resistive), resistive, FLR
  - DW
  - KH
- 3 different eigenfunction regimes
  - Q = quasi-local
  - N = non-local
  - B = barrier-limited non-local
- 3 different saturation/mitigation regimes
  - W = wave-breaking
  - R = Reynold's driven flows
  - M = mean flows
- 2 types of transport
  - C = compact (normally associated with Q)
  - D = distributed (normally associated with N or B)
- Some combinations are more physically interesting: concentrate on what we have seen in past and ongoing SOLT simulations.

## An example: ideal curvature modes in the BWD case

- BWD = barrier-limited, wave-breaking, distributed
- same estimates apply for RX<sup>8</sup> modes
- Starting from  $\lambda_q$  on p. 3,
  - first use  $\tilde{v}_x / \omega = 1/k_x = L_x / \pi$
  - then use  $\gamma = c_s / (R\lambda_p)^{1/2}$
  - to get

$$\lambda_q = \frac{q}{g} \frac{R^{1/2} L_x^2}{\pi^2 \lambda_p^{3/2}}$$

- for compact modes we would set  $\lambda_p = \lambda_q$  and solve for  $\lambda_q$

[8] RX mode are moderate  $k_y \ll k_\eta$  ideal modes in the OM but become resistive (disconnecting from good curvature) near the X-pts due to strong magnetic shear. The growth rate is of order of ideal MHD [J.R. Myra et al., Phys. Plasmas 7, 4622 (2000)]



## Turbulent suppression in H-mode

- This case is QMC in our keyed notation

$$\frac{c_s}{(R\lambda_p)^{1/2}} = \gamma \leq v'_{di} \equiv \frac{c_s \rho_s}{\lambda_p^2}$$
$$\Rightarrow \lambda_p \leq R^{1/3} \rho_s^{2/3}$$

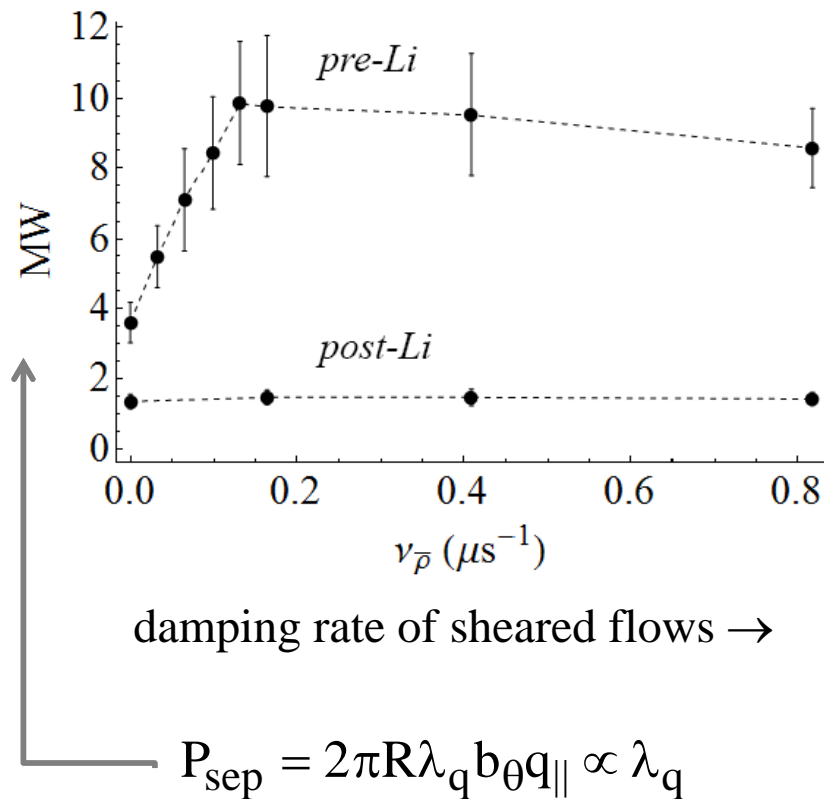
- It is the condition for mean flows to stabilize interchange-like modes
- It provides a rough limit on the pressure gradient in an H-mode assuming  $v_E = -v_{di}$  i.e. that net fluid flows are small.
- It is order of magnitude correct for NSTX
- Even when this condition is satisfied, there can still be instabilities:
  - near the maximum logarithmic pressure gradient the  $E \times B$  shear is zero (assuming  $v_E = -v_{di}$ )
  - low  $k$  non-local distributed modes (N or B and D type) can grow centered at this location and still control the SOL width
  - also curvature enhanced KH modes

## Summary table of some interesting cases

simplified for  $q = g = 5$  (e.g. sheath transmission) to highlight  $R$  and  $\rho_s$  scalings

Instability	Key	Scaling for $\lambda_q$	Remarks
resistive ( $k_\eta$ )	QWC	$2.5 \left( \frac{v_e}{\Omega_e} \right)^{2/7} R^{5/7} \rho_s^{2/7}$	Halpern <sup>3</sup>
ideal or RX ( $\omega_{*i}$ )	QWC	$(R\rho_s)^{1/2}$	larger for low k: NWC $\lambda_q \rightarrow R$ !
ideal or RX	BWC	$R^{1/5} \left( \frac{L_x}{\pi} \right)^{4/5}$	L mode?
drift	QWC	$R^{1/3} \rho_s^{2/3}$	maximal estimate
ideal or RX	_MC	$R^{1/3} \rho_s^{2/3}$	H-mode; mean flow suppression
ideal or RX	NWD	$(\lambda_p R)^{1/2}$	large! even larger if compact; low-k ideal modes destroy SOL
ideal, RX or KH	BRD	$\frac{v}{\Omega_i} \frac{R L_x^2}{\pi^2 \lambda_p \rho_s}$	$\sim v/\lambda_p$ like SOLT; independent of $\gamma$
ideal or RX	BWD	$\frac{R^{1/2} L_x^2}{\pi^2 \lambda_p^{3/2}}$	upper limit (BRD < BWD) SOLT with flow damping?
KH	BWD	$0.2 \frac{L_x^2}{\pi^2} \frac{R \rho_s}{\lambda_p^3}$	
Heuristic Drift		$q\rho_s$	Goldston model <sup>6</sup> ; not instability-based ( $q\rho_s$ is a simplified order of mag. version)

## SOLT simulations<sup>1</sup> for NSTX H-mode suggest BRD, BWD scaling



- linear increase with  $\nu$  in R regime;

$$\lambda_q = \frac{q}{g} \frac{\nu}{\Omega_i} \frac{R L_x^2}{\lambda_p \rho_s}$$

- plateau in W regime

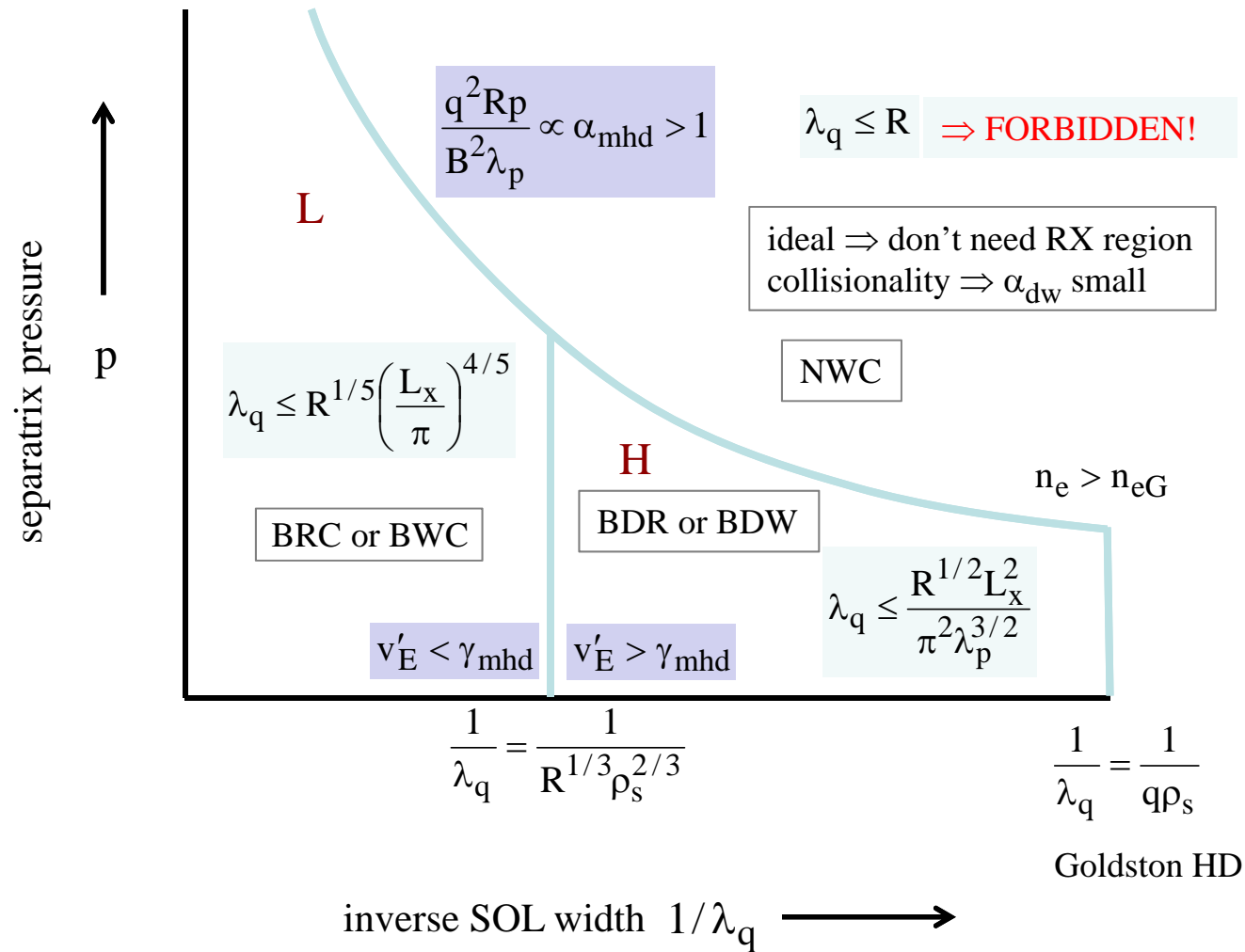
$$\lambda_q = f \frac{q}{g} \frac{R^{1/2} L_x^2}{\lambda_p^{3/2}}$$

- inverse scaling with  $\lambda_p$   
consistent with pre-Li and post-Li [Russell talk]

# Sample SOL width diagram

(with speculative connection to HD density limit)

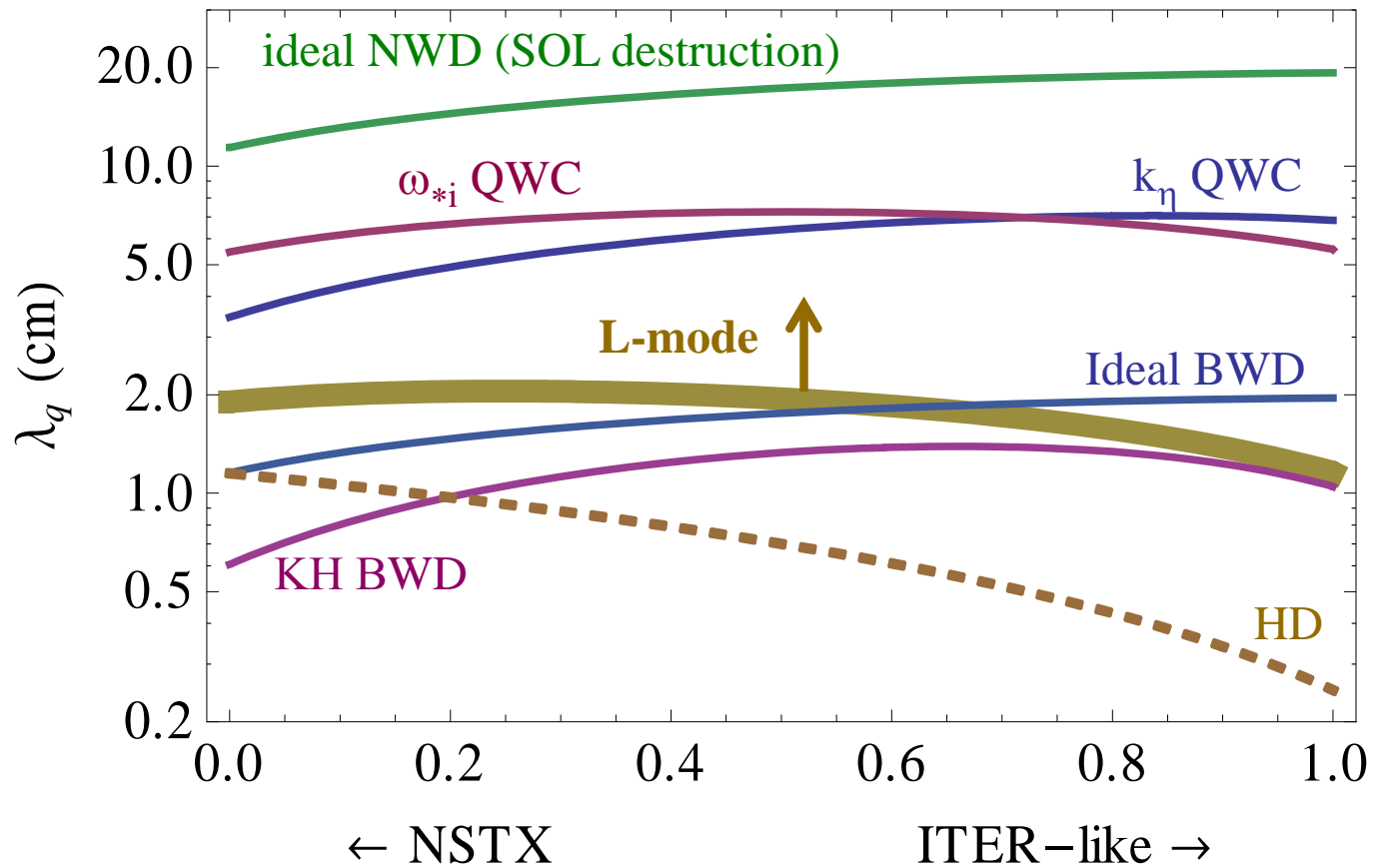
- a connection of heuristic drift (HD) model to density limit was proposed in [Goldston and Eich, 24th IAEA FEC, San Diego, October 8 - 13, 2012, paper IAEA-CN-197/TH/P4-19]



## Notes on the diagram

- This is *not* a regime diagram for the L-H transition; it is meant to show how the predictions for  $\lambda_q$  change in the different regimes.
- L-mode scale lengths  $\lambda_p$  are long, and below the threshold for mean flow suppression. Compact modes are possible.
- The  $R^{1/3} \rho^{2/3}$  boundary only applies to the L-mode side (compact). It gives a relatively wide SOL.
- In H-mode, not only is  $\lambda_p$  shorter, but  $\lambda_q$  is at a different location (distributed) so the resulting SOL width is much narrower than in L-mode, since  $\lambda_p > \lambda_q$ .
- Quoted estimates for  $\lambda_q$  are wave-breaking limit; Reynolds estimates will be smaller.
- When turbulence SOL widths exceed Goldston HD, could get a two-scale SOL; when Goldston HD width is larger, turbulence may be irrelevant.
- Approaching the  $\alpha_{\text{mhd}}$  boundary in H-mode  $\Rightarrow$  increased transport, broadened SOL which moves one up and along the curve to the L-mode regime.
- Strong perpendicular transport at the  $\alpha_{\text{mhd}}$  boundary is consistent with parallel disconnection from the sheaths.

## Order-of-magnitude estimates



- ad-hoc transition of parameters from NSTX ( $x = 0$ ) to ITER-like ( $x = 1$ )
- wave-breaking estimate illustrated; Reynolds estimates will be smaller.
- turbulence results scale better than HD in going to ITER

## Conclusions

- Simple hand-waving considerations for turbulent transport fluxes in various regimes can qualitatively explain some of the SOL width results seen in SOLT simulations: both scaling and order of magnitude.
- Detailed comparison with experiments remains, but present results do not seem unreasonable.
- The turbulent SOL heat flux width in L-mode and H-mode may depend on different transport mechanisms, i.e. separation of driving gradients (pedestal) and responding gradients (SOL) (i.e. compact vs. distributed)
- A speculative relationship is suggested between the turbulence and the heuristic drift mechanism for the SOL width, which may also relate to the density limit.
- Turbulence mechanisms tend to give  $\lambda_q$  a positive scaling with R. These are more favorable for large machines (like ITER) than the HD model which just depends on  $\rho_s$ .